

# Power spectral density

The power spectral density  $S_{xx}$  is the Fourier transform of the autocorrelation  $R_{xx}$  of a [wide-sense stationary](#) process  $x$ .

For CT process  $x(t)$ :

$$S_{xx}(j\omega) \leftrightarrow R_{xx}(\tau)$$

For DT process  $x[n]$ :

$$S_{xx}(e^{j\Omega}) \leftrightarrow R_{xx}[m]$$

## Instantaneous power

Instantaneous power is defined:

$$x^2(t)$$

The expectation of instant power is the autocorrelation with zero time shift:

$$E[x^2(t)] = R_{xx}(0)$$

The expectation of instantaneous power can be written in terms of the power spectral density:

$$E[x^2(t)] = R_{xx}(0) = \frac{1}{2\pi} \int_{-\infty}^{\infty} S_{xx}(j\omega) d\omega$$

Therefore,  $S_{xx}$  describes how instantaneous power is distributed across frequency.

## PSD of filtered process

Consider a process  $y$ , which is the WSS process  $x$  filtered by a function  $h$ :

$$y(t) = (h \ast x)(t)$$

Then, the PSD of this new process is:

$$S_{yy}(j\omega) = |H(j\omega)|^2 S_{xx}(j\omega)$$

## Fluctuation spectral density

Fluctuation spectral density is the power spectral density of the fluctuation of a process from its mean.

In other words, it is the Fourier transform of autocovariance.

$$C_{xx}[m] \rightarrow D_{xx}(e^{j\Omega}) \quad C_{xx}(\tau) \rightarrow D_{xx}(j\omega)$$

## White process

A white process has a flat power spectral density. For a white process  $x(t)$ :

$$S_{xx}(j\omega) = k, \quad -\infty < \omega < \infty$$

## Energy spectral density

Let  $x(t)$  be a random process. Window this signal between  $-T$  and  $T$  to obtain  $x_T(t)$ .  $x_T(t)$  can also be written as:

$$x_T(t) = w_T(t)x(t)$$

where  $w_T(t) = 1$  for  $|t| < T$  and  $0$  otherwise.

The energy spectral density (ESD) is the square of the Fourier transform of the windowed signal  $x_T(t)$ :

$$|X_T(j\omega)|^2$$

The ESD has units “energy/Hz.”

## Periodogram

The periodogram is defined by the ESD divided by the time interval  $2T$ :

$$\frac{1}{2T} |X_T(j\omega)|^2$$

The periodogram has units “power/Hz.”

The limit of the expectation of the periodogram as  $T \rightarrow \infty$  is the power spectral density:

$$S_{xx}(j\omega) = \lim_{T \rightarrow \infty} \frac{1}{2T} E[|X_T(j\omega)|^2]$$

This result is the Einstein-Wiener-Khinchin theorem.

# Spectral estimation

Spectral estimation is estimating power spectral density  $S_{xx}(j\omega)$  or cross spectral density  $S_{xy}(j\omega)$  from experimental or simulated data.

To do this, we replace the expectation  $E[|X_T(j\omega)|^2]$  in the previous section with the average over many iterations from experiments or simulations.

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Last update: **2024-04-30 04:03**